

# SOLAR INDUCED CHLOROPHYLL FLUORESCENCE

OCO-2 LITE FILES (B7000) USER GUIDE

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## Overview Information

The solar induced fluorescence (SIF) OCO-2 Lite files contain a subset of the information in the IMAP-DOAS (IDP) pre-processing L2 files. There is one file per day, for each day that had at least one retrieved sounding. The main purpose of the SIF-lite files is to a) perform post-processing on the original IDP files and b) provide all valid data in significantly smaller files that still contain all necessary information for typical science analyses. In addition, they have some value added:

- Only contain converged soundings that passed initial quality criteria thresholds
- They include an important SIF correction procedure, performed on a daily basis using non-fluorescing surfaces such as deserts or most oceans
- They contain additional information merged from both the A-band preprocessor (ABP) as well as meteorological input data interpolated to the OCO-2 footprint in time and space

The SIF-lite files are provided in the netCDF-4 format ([Unidata hyperlink here](#)). Because netCDF-4 is built on the HDF-5 storage layer, the files may be read with both netCDF4 and HDF5 software, usually available in most computing tools such as Matlab, IDL, python. For python, the use of h5py (<http://www.h5py.org/>) or netCDF4 (<https://github.com/Unidata/netcdf4-python>) is recommended and easily available through most python installations on both Linux/Mac and Windows systems.

Generally speaking, each field in the file is described in the Attributes of that field within the file itself. Descriptions for selected fields are given below, but please be sure to read the Attributes of each used field within the actual Lite file. HDFview ([Hyperlink](#)) is a good tool to quickly browse individual datasets for their content, attributes and so forth. Figure 1 shows a HDFview screenshot, enabling a quick overview of the general structure of the SIF-lite files. Ancillary information from the OCO-2 cloud pre-processors (IMAP-DOAS IDP and Oxygen A-band surface pressure fit ABP) are stored in a separate Group ("Cloud") and information extracted and computed from ECMWF fields interpolated to the OCO-2 footprint are located in the Meteo group. Please note that the ECMWF data is based on the 3hourly forecast system and computation of added value products such as vapor pressure deficit has not been fully validated.

All important variables in the SIF-lite file should have corresponding attributes for units and standard names as well as a somewhat longer comment to describe the fields in more detail.

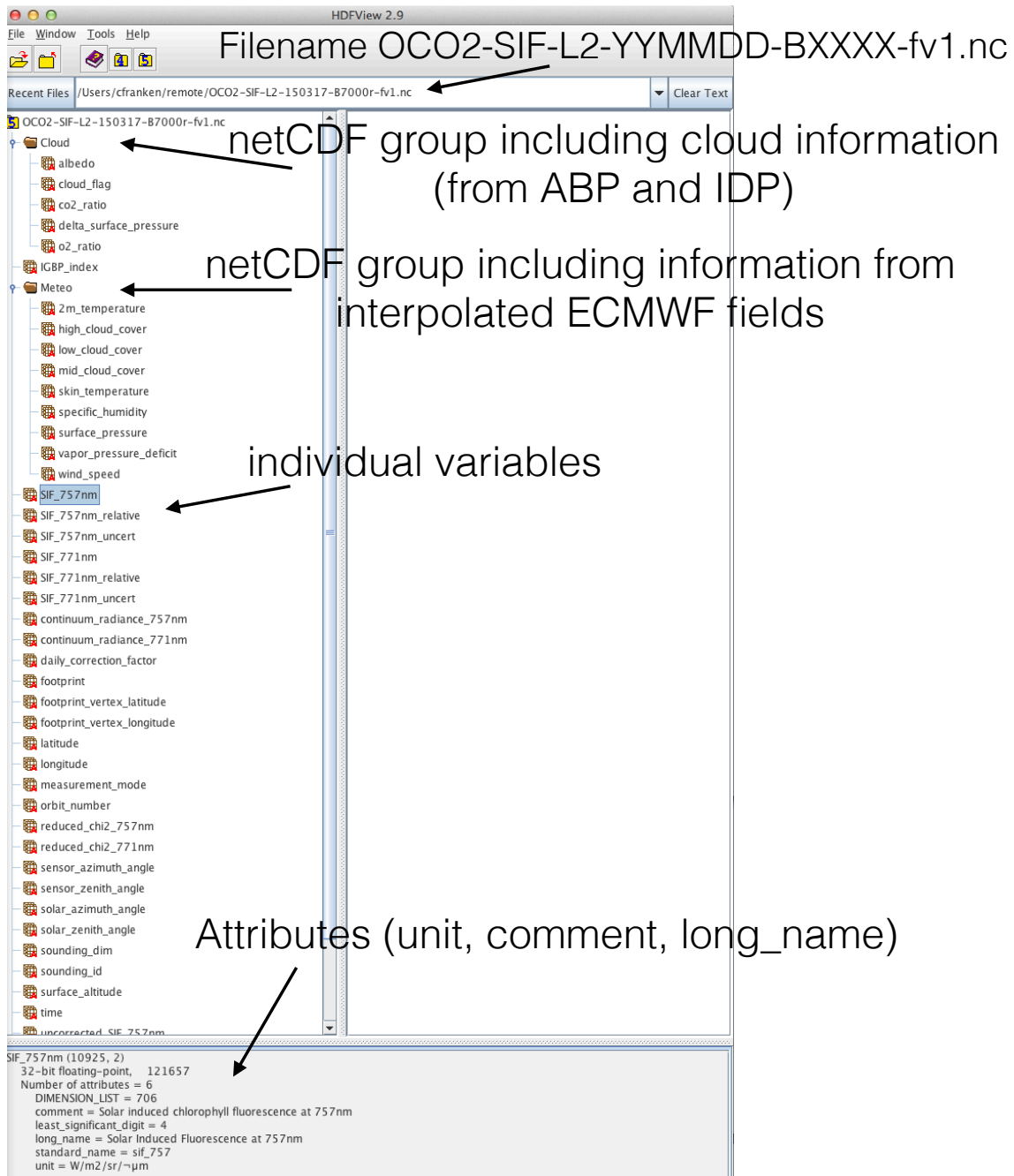


Figure 1: HDFview screenshot outlining the content of a SIF-lite file.

## File Structure & Fields

The primary variables that most users will need exist at the main level. In addition, there are some additional variables that certain users might want, contained in two groups within the file: Cloud and Meteo. Some NetCDF readers may not see these groups; if this happens, please update your NetCDF reader or use an HDF-5 reader.

### Main level variables

Below, we will summarize all variables on the main level, starting with the most important variable first (those which initial users will focus on in the beginning).

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#### KEY VARIABLES

**latitude** Latitude of the center of the OCO-2 footprint in the WGS 84 reference frame (degrees)

**longitude** Longitude of the center of the OCO-2 footprint in the WGS 84 reference frame. (degrees)

**time** time in seconds since 1993-01-01 00:00:00

**SIF\_757nm** Solar Induced Fluorescence at 757nm ( $\text{W}/\text{m}^2/\text{sr}/\mu\text{m}$ )

**SIF\_771nm** Solar Induced Fluorescence at 771nm ( $\text{W}/\text{m}^2/\text{sr}/\mu\text{m}$ ). This value is typically about 1.5 times smaller than at 757nm and the two fields could be averaged after multiplication of `sif.771` with 1.5. However, we suggest to use them independently in the beginning as both may be affected by different bias sources and could provide robustness in the data analysis if both agree well.

**SIF\_757nm\_uncert**  $1-\sigma$  uncertainty in retrieved SIF at 757nm ( $\text{W}/\text{m}^2/\text{sr}/\mu\text{m}$ )

**SIF\_771nm\_uncert**  $1-\sigma$  uncertainty in retrieved SIF at 771nm ( $\text{W}/\text{m}^2/\text{sr}/\mu\text{m}$ ). It is important to not that the  $1\sigma$  errors for both variables can be substantial. This can lead to negative SIF values, which are perfectly valid in a measurement retrieval sense owing to the presence of retrieval noise. Discarding negative values is dangerous as it biases the average. In general, multiple soundings will need to be averaged to reduce the noise by a factor of  $1/\sqrt{n}$ , with  $n$  being the number of averages.

**Cloud/o2\_ratio** Ratio of retrieved vs. predicted  $\text{O}_2$  column using the 771nm retrieval window (used for initial cloud screening, usually values  $>0.9$  are sufficient for SIF related cloud screening as the ABP filters may be too strict). The most cloudy scenes are already pre-filtered in this dataset.

**daily\_correction** Daily Correction Factor Correction factor to estimate daily average SIF from instantaneous SIF (using pure geometric incoming light scaling). Especially at high latitudes, the fluorescence signal at 13:30 local solar overpass time cannot be directly compared with GPP since the length of day and variability of the solar zenith angle has to be taken into account. Under cloud-free conditions and ignoring Rayleigh scattering as well as gas absorption, the downwelling solar radiation scales linearly with  $\cos(SZA)$ . If  $t_0$  denotes the time of measurement in fractional days, a first order approximation for a daily fluorescence average can be written as:

$$\overline{F_S} = F_s / \cos(SZA(t_0)) \cdot \int_{t=t_0-12h}^{t=t_0+12h} \cos(SZA(t)) dt$$

The aforementioned correction  $\cos(SZA(t_0)) \cdot \int_{t=t_0-12h}^{t=t_0+12h} \cos(SZA(t)) dt$  is saved in *daily\_correction*, computing the integral numerically in 10 minute time-steps (Using pyEphem (<http://rhodesmill.org/pyephem/>) to compute SZA as a function of latitude, longitude and time).

**IGBP\_index** IGBP Index One-Minute Land Ecosystem Classification Product is a global (static map) data set of the International Geosphere-Biosphere Programme (IGBP) classification scheme stored on an equal-angle rectangular grid at 1-minute resolution. See <http://modis-atmos.gsfc.nasa.gov/ECOSYSTEM/>

**solar\_zenith\_angle** Solar Zenith Angle (degrees): Solar zenith angle is the angle between the line of sight to the sun and the local vertical

**measurement\_mode** OCO-2 Measurement mode, 0=Nadir, 1=Glint, 2=Target OCO-2 Measurement mode, 0=Nadir, 1=Glint, 2=Target, users should separate those for the analysis!

#### ADDITIONAL GEOMETRY/TIME VARIABLES

**sensor\_zenith\_angle** Sensor Zenith Angle (degrees): Sensor zenith angle is the angle between the line of sight to the sensor and the local vertical

**sensor\_azimuth** Sensor Azimuth Angle (degrees): Azimuth angle between line of sight and local north

**solar\_azimuth** Solar Azimuth Angle (degrees): Azimuth angle between the solar direction as defined by the sounding location, and the sounding local north

**time** time in seconds since 1993-01-01 00:00:00

**footprint\_id** Detector Footprint Number: OCO-2 footprint identifier (1-8), identifying the 8 independent OCO-2 spatial samples per frame

**sounding\_id** Unique Identifier for each sounding (YYYYMMDDHHMMSS)

**altitude** Surface Altitude Surface altitude of observed footprint (meters)

**orbit\_number** Orbit Number

**footprint\_latitude\_vertices** Latitude corner coordinates of the sounding location

**footprint\_longitude\_vertices** Longitude corner coordinates of the sounding location

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#### OTHER VARIABLES

**uncorrected\_SIF\_771nm** raw Solar Induced Fluorescence at 771nm (without any bias correction) ( $\text{W}/\text{m}^2/\text{sr}/\mu\text{m}$ ) Solar induced chlorophyll fluorescence at 771nm (without any bias correction)

**uncorrected\_SIF\_757nm** raw Solar Induced Fluorescence at 757nm (without any bias correction) ( $\text{W}/\text{m}^2/\text{sr}/\mu\text{m}$ ) Solar induced chlorophyll fluorescence at 757nm (without any bias correction)

**uncorrected\_SIF\_771nm\_relative** Solar Induced Fluorescence at 771nm in fractions of continuum level Solar induced chlorophyll fluorescence at 771nm in fractions of continuum level

**uncorrected\_SIF\_757nm\_relative** Solar Induced Fluorescence at 757nm in fractions of continuum level, no bias correction Solar induced chlorophyll fluorescence at 757nm in fractions of continuum level, no bias correction

**continuum\_radiance\_757nm** Continuum level radiance in the 757nm retrieval window ( $\text{W}/\text{m}^2/\text{sr}/\mu\text{m}$ )

**continuum\_radiance\_771nm** Continuum level radiance in the 771nm retrieval window ( $\text{W}/\text{m}^2/\text{sr}/\mu\text{m}$ ), note: this multiplies the OCO-2 radiance with a factor 2 to account for unpolarized light and units are converted as well.

**reduced\_chi2\_757nm** reduced  $\chi^2$  of the retrieval residuals in the 757nm window

**reduced\_chi2\_771m** reduced  $\chi^2$  of the retrieval residuals in the 771nm window

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## Meteo group variables

- Specific\_humidity** Specific humidity at surface layer at the sounding location (ECMWF forecast)
- vapor\_pressure\_deficit** Vapor pressure deficit at the sounding location (2m) (ECMWF forecast, Pa)
- skin\_temperature** K Skin temperature at the sounding location (ECMWF forecast, K)
- 2m\_temperature** 2m temperature at the sounding location (ECMWF forecast)
- wind\_speed** surface wind speed at sounding location (ECMWF forecast, m/s)
- low\_cloud\_cover** Cloud cover for low clouds (ECMWF forecast)
- mid\_cloud\_cover** Cloud cover for medium alt. cloud (ECMWF forecast)
- high\_cloud\_cover** Cloud cover for high clouds (ECMWF forecast)
- surface\_pressure** Surface pressure at the sounding location (ECMWF forecast, Pa)
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## Cloud group variables

- surface\_albedo\_760** Surface albedo (lambertian equivalent) as retrieved in the ABO2 preprocessor at 760nm
- cloud\_flag** Cloud Flag from ABO2 (0-Classified clear, 1=Processing failed, 2=Not classified)
- delta\_pressure** Retrieved-predicted surface pressure (Pa) from ABO2 (usable as cloud screener)
- co2\_ratio** Ratio of CO<sub>2</sub> retrieved in weak and strong CO<sub>2</sub> band (value near 1 indicate scattering free scene)
- o2\_ratio** Ratio of retrieved vs. predicted O<sub>2</sub> column using the 771nm retrieval window (used for initial cloud screening, ususally values >0.9 are sufficient for SIF related cloud screening as the ABP filters may be too strict). The most cloudy scenes are already pre-filtered in this dataset.

## Quality Filtering

The data has been pre-filtered using the following criteria:

**altitude**  $> 0$  to exclude oceans

**SIF** between  $-5$  and  $50 \text{ W/m}^2/\text{sr}/\mu\text{m}$  to exclude obvious high and low fliers (rare)

$\chi^2$  between  $0$  and  $2$  to exclude fits that don't represent the spectrum well

**Continuum Level Radiance** between  $28$  and  $195 \text{ W/m}^2/\text{sr}/\mu\text{m}$  (too low and too bright scenes might be more affected by detector calibration)

**Solar Zenith Angle**  $< 70$  to remove retrievals with too high airmasses, for which rotational Raman scattering might be important

**O2 ratio** between  $0.85$  and  $1.5$  to exclude most cloudy scenes

**CO2 ratio** between  $0.5$  and  $4$  to exclude most cloudy scenes

It has to be noted that this quality filtering is much relaxed compared to the OCO-2 XCO<sub>2</sub> lite file generation as SIF is less susceptible to atmospheric scattering. In addition, SIF retrievals are, as opposed to XCO<sub>2</sub>, accurate but imprecise, so that averaging of multiple data-points is required. Also note that clouds don't affect SIF signal strongly (Frankenberg et al, 2012).

## Bias correction

The SIF retrievals presented here are based on the retrieval methodologies described in Frankenberg et al (2011a,b). The main retrieval quantity in the retrieval state vector is relative fluorescence, i.e. the fractional contribution of SIF to the continuum level radiance. Absolute SIF is being generated in the post-processing step. Owing to various effects such as uncertainties in the exact instrument line-shape per footprint or slight uncertainties in detector linearity, biases in retrieved SIF can occur. Here, we follow a similar strategy as in Frankenberg et al (2011b), using reference targets to correct for biases in SIF. Owing to the large amount of OCO-2 data, we can now perform these corrections on a daily basis. An example is shown in Figure 2.

There is room for improvement for this bias correction, specifically treating it separately for Glint and Nadir modes and using temporal smoothing of the correction curves. We appreciate any feedback as to the data quality and consistency to keep improving OCO-2 SIF-lite generation. The raw relative and absolute SIF values are provided in the datasets as well (see description above).



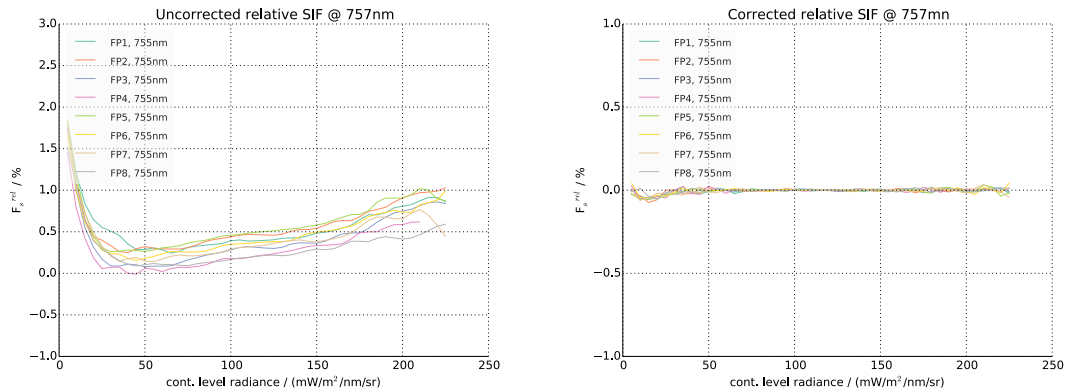


Figure 2: SIF-lite bias correction as a function of signal level (example taken from Oct. 2, 2014). Left: raw relative fluorescence retrievals over supposedly non-fluorescing surfaces. Right: Dataset after subtraction of the biases determined as a function of signal level

## References

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